# The influence of explicit local dynamics on range expansions driven by long-range dispersal

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Range expansions are common in natural populations. They can take such forms as an invasive species spreading into a new habitat or a virus spreading from host to host during a pandemic. When the expanding species is capable of dispersing offspring over long distances, population growth is driven by rare but consequential long-range dispersal events that seed satellite colonies far from the densely occupied core of the population. These satellites accelerate growth by accessing unoccupied territory, and also act as reservoirs for maintaining neutral genetic variation present in the originating population, which would ordinarily be lost to drift. Prior theoretical studies of dispersal-driven expansions have shown that the sequential establishment of satellites causes initial genetic diversity to be either lost or maintained to a level determined by the breadth of the distribution of dispersal distances. If the tail of the distribution falls off faster than a critical threshold, diversity is steadily eroded over time; by contrast, broader distributions with a slower falloff allow some initial diversity to be maintained for arbitrarily long times. However, these studies used lattice-based models and assumed an instantaneous saturation of the local carrying capacity after the arrival of a founder. Real-world populations expand in continuous space with complex local dynamics, which potentially allow multiple pioneers to arrive and establish within the same local region. Here, we evaluate the impact of local dynamics on the population growth and the evolution of neutral diversity using a computational model of range expansions with long-range dispersal in continuous space, with explicit local dynamics that can be controlled by altering the mix of local and long-range dispersal events. We found that many qualitative features of population growth and neutral genetic diversity observed in lattice-based models are preserved under more complex local dynamics, but quantitative aspects such as the rate of population growth, the level of maintained diversity, and the rate of decay of diversity all depend strongly on the local dynamics. Besides identifying situations in which modeling the explicit local population dynamics becomes necessary to understand the population structure of jump-driven range expansions, our results show that local dynamics affects different features of the population in distinct ways, and can be more or less consequential depending on the degree and form of long-range dispersal as well as the scale at which the population structure is measured.

## I. INTRODUCTION

Range expansion—the act of a population expanding 6 7 into new territory—is common in biological populations. Range expansions occur naturally and randomly all the 8 time, often as the result of a species' natural movement, 9 such as by animals moving into new territory or maple 10 helicopters carrying seeds away from their parent tree. 11 Researchers have documented range expansions in a wide 12 variety of organisms, such as plants [1], birds [2], sea 13 creatures [3, 4], and terrestrial animals [5, 6], even hu-14 mans [7]. Range expansions are increasingly forced by 15 global warming as the changing climate makes traditional 16 habitats inhospitable, while potentially opening up new 17 hospitable regions [8]. 18

Range expansions leave distinctive signatures in the patterns of genetic diversity of a population that can mimic the effects of natural selection [9]. Individuals at the frontier of an expanding population make a large contribution to the subsequent expansion wave, even if their frontier position was solely due to chance; as a result, genetic variants they carry can acquire high fre-

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<sup>26</sup> quencies in the population in a phenomenon termed gene <sup>27</sup> surfing [10, 11]. Independent surfing events in separate <sup>28</sup> sections of the expansion front cause the population to <sup>29</sup> segregate into genetically distinct sectors, promoting an <sup>30</sup> illusion of local adaptation from purely neutral muta-<sup>31</sup> tions [12–15]. Modeling the combined effect of spatial <sup>32</sup> structure and stochasticity on neutral genetic diversity <sup>33</sup> is key to understanding the biological origins of estab-<sup>34</sup> lished genetic patterns, and to the successful prediction <sup>35</sup> of future genetic diversity in pandemics and ecological <sup>36</sup> expansions.

37 The influence of random chance on genetic diversity <sup>38</sup> during range expansions can be amplified by long-range <sup>39</sup> dispersal [16]. Many species have evolved ingenious ways <sup>40</sup> of dispersing offspring over long distances with help from <sup>41</sup> natural forces and from other organisms [17]. Plants rely <sup>42</sup> on the dispersal of seeds and pollen by wind, waves, and <sup>43</sup> animals [18]. Glacier ice worms can travel hundreds of <sup>44</sup> miles, likely carried by migratory birds [19]. Modern pan-<sup>45</sup> demics are driven by microorganisms hitchhiking on air <sup>46</sup> travelers to find new uninfected populations [20]. Even <sup>47</sup> if long-range dispersal events are rare, they have an out-<sup>48</sup> sized influence on the expansion because they enable pi-<sup>49</sup> oneers to seed satellite colonies in uninhabited areas. If <sup>50</sup> a pioneer happens to land in a place with abundant re-<sup>51</sup> sources and little to no competition, its descendants may <sup>52</sup> flourish. The pioneer's genes will then propagate and any

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54 55 56 57 58 59 60 chance of contributing to growth. The evolution of over- <sup>119</sup> to the formation of sectors [12, 30]. 61 all diversity during the range expansion is governed by 120 62 63 64 by the population [26, 27]. 65

66 67 68 69 70 71 72 73 74 75 76 77 78 79 is exclusively short-range [28]. 80

81 82 description of the dispersal behavior of many species [29], <sup>141</sup> approximate a continuum [35]. 83 power-law kernels are a useful tool for analyzing and clas- 142 84 85 to long-range dispersal [16]. The exponent  $\mu$  is a key fac- 144 saturation and population management. 86 87 89 90 91  $\mu = 0$  to ensure a normalizable distance distribution). 150 rate [36, 37] or let the growth be determined by random 93 94 95 97 98 extremes in a tunable manner. 99

100 <sup>101</sup> of population growth dynamics [16] and spatial genomic <sup>159</sup> to the next [36, 37]. Since most computational studies 102 103 104 105 of the kernel exponent  $\mu$  below which the population  ${}^{_{164}}$  open question. 106 grows nearly as fast as a well-mixed population, and a 165 107 <sup>108</sup> significant fraction of the neutral variation in the orig-<sup>166</sup> ics and neutral evolution of range expansions driven by <sup>109</sup> inating population is preserved for arbitrary long times <sup>167</sup> power-law kernels [16, 27] were derived using a lattice of <sup>110</sup> due to serial reintroduction of variants from the core of <sup>168</sup> demes with an additional simplifying assumption: upon

<sup>53</sup> genetic variants they carry will reach high frequencies in <sup>111</sup> the expanding population. For kernel exponents close the vicinity of the satellite [21-25] even in the absence of  $_{112}$  to but above the critical threshold, population growth a selective advantage; random chance alone has caused 113 is slowed down dramatically and neutral diversity is the pioneer's genes to become prominent by means of a  $_{114}$  steadily eroded. However, at even higher values of  $\mu$ , the founder effect, leading to a *suppression* of local diversity <sup>115</sup> behavior approaches that of short-ranged jump kernels, within satellites. However, long-range dispersal also fa- 116 where the population advances as a front moving outward vors neutral diversity at larger scales, by ensuring that 117 at constant speed. In this situation, a small fraction of individuals well within the expanding population have a 118 the diversity in the originating population persists due

Less well understood is the influence of the second key the trade-off between the two effects, and can depend sen- 121 feature of spatial population models: the density regulasitively on the degree of long-range dispersal experienced 122 tion mechanism. Modeling growing populations in a spa-123 tial continuum presents challenges to both the forward-Modeling the general characteristics of range expan-<sup>124</sup> in-time [31] and backward-in-time [32, 33] approaches, sions requires two minimal ingredients: a probability dis- 125 due to the necessity of systematically imposing a local retribution of dispersal distances J(r), also called the jump  $_{126}$  gion of influence within which each individual can impact kernel, from which dispersal events are randomly drawn; 127 the growth of its neighbors. Local density regulation is and a method of local density regulation to model the 128 commonly implemented by dividing up space into a reguexistence of a finite carrying capacity. When long-range 129 lar grid of well-mixed subpopulations called demes, each jumps are present, the tail of the jump kernel, i.e. its 130 of which has a fixed carrying capacity. Migration events, behavior at long distances, critically influences the fate 131 drawn from the jump kernel, transport individuals across of the population at long times. Fundamental differences 132 demes. Deme-based models and their variants are widely from short-range dispersal are observed when the jump 133 used in population genetics [34], including for the study kernel is "fat-tailed"; i.e. it decays slower than exponen- 134 of range expansions [9, 30]. However, models that rely tially with increasing distance. Fat-tailed jump kernels 135 on a lattice of demes have their limitations. By design, lead to expansions that accelerate as they progress, unlike 136 they do not capture spatial structure and stochasticity the constant-speed expansions that occur when dispersal <sup>137</sup> at scales smaller than the effective deme size. Impos-138 ing an artificial grid of demes also introduces artifacts A commonly used fat-tailed kernel is the power-law <sup>139</sup> to the population structure, which can in some instances jump kernel  $J(r) \sim 1/r^{\mu+1}$ . Besides providing a good <sup>140</sup> get worse upon increasing the grid resolution to better

Additionally, using deme-based models forces resifying the breadth of potential population outcomes due 143 searchers to make decisions about the specifics of deme The followtor governing the long-time characteristics of the growth 145 ing selection of recent work exemplifies various possible and the dispersal patterns, whereas other details of J(r) 146 strategies. Some may choose to have demes that instansuch as its short-distance functional behavior are less rel- 147 taneously change from being empty to full upon the arevant [16]. A broad range of expansion behaviors is en- 148 rival of the first migrant [16, 27], while others may let compassed by varying the kernel exponent (limited to 149 the deme population grow logistically at a predetermined At high  $\mu$ , the jump kernel decays quickly with increas- <sup>151</sup> migration events that bring in individuals from other ing distance, and a colony expands at a constant rate as 152 demes [38]. Death can occur in various ways, such as by if there were exclusively short-range dispersal. As  $\mu \to 0$ , 153 attempting to disperse into an already full deme [16, 27] spatial structure becomes irrelevant and a colony grows <sup>154</sup> or by being randomly resampled out of an overfull deme's as if it were in a well-mixed liquid environment. The 155 population [38]. If the density regulation unit is the deme intermediate range of kernel values connects these two 156 population as a whole rather than the individuals in the <sup>157</sup> deme, death may not explicitly occur to any individuals, Recent work has catalogued the distinctive features <sup>158</sup> but the deme population size changes from one time step patterns [27] that can be achieved upon varying the ker- 160 involving long-range dispersal, including the quoted prior nel exponent in range expansions driven by power-law <sup>161</sup> results [16, 21-25, 27, 36, 37, 39], have relied on demegrowth kernels (a detailed summary is provided in Sec- 162 based approximations, the applicability of their conclution II). These studies have identified a critical value 163 sions to continuum-space population growth remains an

The aforementioned results on the population dynam-

171 172 174 175 176 177 most offspring land within a short distance of their par- 232 spring according to a Poisson process; offspring attempt 178 179 180 181 182 183 184 186 187 tures left behind by long-range dispersal is unknown. 188

189 190 191 based simulations of range expansions driven by power- 247 as summarized below. 192 law kernels. Our simulations were implemented in the 193 population genetics program SLiM [40], and do not use 194 a grid of demes or assume instant saturation of the local 248 A. Population growth and time-doubling hierarchy 195 carrying capacity by the first arrival. Instead, individ-196 uals occupy positions in continuum space and their sur-197 vival depends on the number of other individuals present 198 within a defined region of influence at the time of their 199 birth. When possible, we compared the outputs to the 200 predictions from models based on lattices of demes of 201 202 the population growth rate [16] and the evolution of neutral genetic diversity [27]—we term these prior models 203 'lattice-based" predictions. We found that our results 204 often agreed with the lattice-based predictions, giving 205 conditional support to prior results based on models that 206 only focus on the founders. However, when individuals 207 can share resources with many others, we found that fo-208 cusing exclusively on the founders misses important dy-209 <sup>210</sup> namics between coexisting or competing alleles. In those 211 cases, it becomes necessary to also consider individu-<sup>212</sup> als who arrive after the pioneer. We identify parameter <sup>213</sup> regimes where using the lattice-based models is justified, <sup>214</sup> and show that they depend on the specific kernel expo-215 nent.

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### II. BACKGROUND

217 expansion dynamics for populations experiencing long- 273 on  $\mu$  and d). 218 range dispersal with fat-tailed kernels, which were ob- 274 219 220 221 <sup>223</sup> unoccupied space when offspring are dispersed accord- <sup>278</sup> were seeded at an earlier time by a rare but consequen-

169 arrival at an empty deme, the pioneer immediately sat- 224 ing to a power law jump kernel that decays according urates the deme, excluding any other migrants from es-  $_{225}$  to  $J(r) \sim 1/r^{\mu+1}$ . The authors showed that the powertablishing themselves. Not only does this assumption ex- 226 law tail captures the qualitative features of the long-term clude any effects of local dynamics on population growth, 227 population growth, and that the short-range behavior of it also enforces a local *founder takes all* effect where 228 the jump kernel has a negligible impact on the long term only one migrant is allowed to contribute to the genetic 229 population growth. The model of Ref. 16 (hereafter "the makeup of a density regulation region. Instant local sat- 230 lattice model") divides d-dimensional space into a lattice uration is justified when long-range jumps are rare and 231 of habitats or "demes". Occupied demes generate offents; then, the local logistic growth within a deme oc- 233 to migrate to a new deme randomly chosen by drawing curs extremely fast compared to the typical time to ar-  $^{234}$  a dispersal distance from J(r) and a random direction rival of another migrant from a different deme, and can 235 relative to the originating deme. Instant local saturabe treated as instantaneous. However, the instant satu- 236 tion is assumed and is enforced in the model by allowing ration and founder-takes-all assumptions can be invalid 237 only two states to each deme: occupied, or empty. A when the time scales of local and long-range dispersal 238 migration attempt to an empty deme is successful, and are comparable, in which case a local region might receive 239 immediately turns the state of that deme to occupied. and send out several migrants while it is being saturated. 240 A migration attempt to an occupied deme is unsuccess-The influence of the breakdown of fast local saturation on 241 ful, and the offspring dies. These assumptions guarantee the population dynamics and the spatial genomic struc- 242 a founder-takes-all effect at the local level. Henceforth, <sup>243</sup> when we refer to the lattice model, it is implied that In this work, we address these gaps in our knowledge 244 instant local saturation and local founder-takes-all are of range expansions driven by long-range dispersal by 245 enforced. Much of our current understanding of jumpperforming and analyzing continuum space, individual- 246 driven range expansions derives from the lattice model,

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Analysis of the lattice model [16] showed that at all 240  $_{250}$  times t, a core region of the colony can be identified that <sup>251</sup> is centered at the originating population of the range ex-<sup>252</sup> pansion and within which most demes are occupied. The 253 size of this core region is proportional to the total pop-<sup>254</sup> ulation size M(t) of occupied demes in the expansion. <sup>255</sup> The long-time asymptotic behavior of the radius of the <sup>256</sup> core region  $\ell(t) \propto [M(t)]^{1/d}$  depends on the "heaviness" <sup>257</sup> of the tail of the jump kernel, which is set by the kernel <sup>258</sup> exponent  $\mu$ . There are two distinct growth possibilities, 259 separated by the value  $\mu = d + 1$ : the colony expands  $_{260}$  at a constant rate for  $\mu > d+1$  and it expands faster <sup>261</sup> than linearly when  $\mu < d+1$  [16]. The faster-than-linear <sup>262</sup> growth regime is driven by long jumps whose character-263 istic size continues to increase as the core expands: this 264 "jump-driven" growth regime, in which pioneers have a 265 large impact on growth, will be the focus of this pa-<sup>266</sup> per. Within the jump-driven regime, a second special 267 value  $\mu = d$  separates two distinct asymptotic behav-<sup>268</sup> iors of the core growth at long times  $(t \to \infty)$ : when  $_{269} d < \mu < d+1$ , the core grows asymptotically as a power <sup>270</sup> law which is faster-than-linear in time  $(\ell(t) \propto t^{\frac{1}{\mu-d}})$ , in  $_{271}$  contrast to stretched-exponential growth when  $\mu < d$ We first summarize prior results [16, 27] on range  $_{272}(\ell(t) \propto \exp(B_{\mu}t^{\eta}))$ , where  $B_{\mu}$  and  $\eta$  themselves depend

A key result of Ref. 16 was that expansions in the tained using lattice models conforming to the founder- 275 jump-driven regime are governed by a hierarchical timetakes-all assumption at the deme level. Ref. 16 used 276 doubling structure, as depicted in Fig. 1. The core of a lattice model to quantify how a colony expands into 277 the colony expands by "absorbing" satellite colonies that 303



FIG. 1. Schematic diagram of the time doubling hierarchy discovered by Ref. 16. Shaded parts of the plot represent regions of space that are occupied at a given time. The core of the colony (central funnel) grows by absorbing satellites that were seeded at an earlier time by long-range dispersal. A typical satellite being absorbed into the core at time T(smaller funnel at right) was seeded at time of order T/2 by order  $\ell(T/2)$  when it merges with, and becomes part of, the core.

<sup>279</sup> tial long jump. A typical satellite being absorbed into the 280 core at time t was seeded approximately at time t/2 by <sup>281</sup> an offspring who dispersed roughly a distance  $\ell(t)$  from 282 its parent in the core of the colony. Mathematically, this 283 self-consistency condition can be expressed as

$$\ell(t)^{d+\mu} \sim t\ell(t/2)^{2d},$$
 (1)

287 288 290 291 292 293 294 295 296 297 <sup>298</sup> are present in a population experiencing long-range dis-<sup>355</sup> progresses, albeit at a slow rate. As  $\mu$  approaches d, the <sup>299</sup> persal [27, 39]. Note that the time-doubling hierarchy <sup>356</sup> heterozygosity approaches a finite value but the conver-300 only relies on the assumption of instant local saturation, 357 gence to this value becomes extremely slow and cannot <sup>301</sup> and does not require that space be discretized into a reg-<sup>358</sup> be observed over practical simulation times. Notably, for 302 ular lattice of demes.

### в. Persistence of initial neutral variation

A striking consequence of range expansions is that the 304 305 combination of stochasticity and spatial structure can leave behind patterns of neutral genetic variation that are typically associated with selection, such as sweep-like enrichment of individual alleles [10], diversity gradients [11], 308 and segregation of variants into distinct regions [12]. Simplified models of neutral evolution in spatially structured populations enable us to understand such patterns 311 312 and to distinguish them from the outcomes of selective events. One aspect of neutral variation that is closely <sup>314</sup> tied to the mode of dispersal is the persistence of initial <sup>315</sup> genetic diversity in the originating population during its <sup>316</sup> expansion into new territory [21, 26]. When dispersal is 317 exclusively short-ranged, only individuals near the edge 318 of the range expansion contribute to future variation; in the absence of new mutations, much of the initial di-319 320 versity can be lost over time due to successive founder <sup>321</sup> events at the edge. Long-range dispersal enables regions <sup>322</sup> far within the population to contribute to the expansion, <sup>323</sup> which maintains their alleles in the growing population an offspring who dispersed a distance of roughly  $\ell(T)$  from  $_{324}$  and favors diversity. However, founder effects are not its parent in the core of the colony; it has grown to a size of 325 eliminated: each long-range jump seeds a satellite out-<sup>326</sup> break in which all offspring share the allelic identity of 327 the seeding pioneer, acting as a genetic bottleneck which 328 eliminates diversity locally in the absence of mutations. 329 The fate of the initial neutral variation as the expansion <sup>330</sup> progresses is determined by the balance between these contrasting effects.

The evolution of initial neutral diversity in jump-332 driven range expansions was analyzed in Ref. 27, us-333 <sup>334</sup> ing a lattice model in which neutral variation was intro-<sup>335</sup> duced in the starting population, and no new mutations 336 appeared during the expansions. The existence of the <sup>337</sup> time-doubling hierarchy, Eq. (1), was used to identify an 338 effective population of homogeneous satellites whose evo-339 lution captures the balance between diversification and <sub>340</sub> coarsening for a given jump kernel exponent. As with the <sup>284</sup> where the tilde signifies agreement of the leading func-<sup>341</sup> behavior of the core radius growth, the amount of initial tional dependence of either side of the relation on the 342 diversity preserved after a range expansion was shown to time variable, without including time-independent pref- 343 suffer different fates depending on the value of the kernel actors or terms whose fractional contributions vanish at  $_{344}$  exponent relative to the spatial dimension. When  $\mu < d$ . long times. The time-doubling hierarchy and Eq. (1) <sup>345</sup> the diversifying influence of long jumps dominates; note form the basis for deriving the asymptotic functional 346 the large number of satellites well separated from the forms of  $\ell(t)$  summarized above; unlike those asymptotic  $_{347}$  core in Fig. 2g-i. The seeding of many satellites by longforms that are valid only at very long times  $t \to \infty$ , the  $_{348}$  range dispersal events from the core enables the popuself-consistency condition holds as long as the population <sup>349</sup> lation to preserve a finite amount of its initial heterozyis large enough that an appreciable number of long-range  $_{350}$  gosity at long times. By contrast, when  $d < \mu < d + 1$ , jumps have occurred [16]. Equation (1) forms a basis for 351 the local coarsening of diversity due to bottlenecks bemore accurate functional forms of the outbreak growth 352 comes more significant; note the small number of large dynamics [16], and also leads to quantitative insights into  $_{353}$  monoclonal satellites in Fig. 2d-f. The heterozygosity the evolution of genetic diversity when multiple variants 354 decays inexorably towards zero as the range expansion  $_{359} \mu > d+1$ , some diversity is also preserved at long times

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FIG. 2. Snapshots of simulated range expansions at different population sizes M. These simulations began with 100 individuals equally split between two neutral alleles, labeled as either purple or yellow in these plots. a-c.) Diversity is preserved by the formation of monoallelic sectors for  $\mu > d + 1$ . d-f.) The small number of satellite outbreaks act as bottlenecks, eroding diversity for  $d < \mu < d + 1$ . g-i.) Long-range jumps transport alleles from the core to the exterior of the colony, preserving diversity for  $\mu < d$ . Additional parameters are K = 10 and p = 0.

<sub>360</sub> due to the formation of sectors in outward range expan- <sub>384</sub> model which we use to investigate these questions. <sup>361</sup> sions, as shown in Fig. 2a–c [12, 30]. Jump kernels of intermediate breadth  $(d < \mu < d + 1)$  therefore support 362 lower neutral diversity than broader ( $\mu < d$ ) and nar-363 rower  $(\mu \ge d+1)$  kernels. 364

In summary, Ref. 27 established that long-range dis- 386 In order to study jump-driven range expansions which 365 366 368 369 371 can erase even the limited heterozygosity preserved due 392 stant rate, and offspring attempt to establish themselves 372 373 374 375 376 377 378 380 381 <sup>382</sup> plicit local dynamics on neutral genetic variation as the <sup>403</sup> persal and spatial structure during spreading. Second, 383 expansion progresses. We next introduce our simulation 404 once offspring are successfully established, they do not

### III. METHODS

persal can preserve some of the genetic diversity from the 387 rely on neither a lattice nor the assumption of instant looriginating population at long times, but only for jump 388 cal dynamics, we used the evolutionary simulation softkernels broader than a dimension-dependent threshold. 389 ware SLiM [40] to simulate range expansions on a 2D Narrower kernels cause diversity to erode over the course 390 continuous landscape without restricting ourselves to a of the expansion due to successive founder events, which 391 lattice of demes. Individuals produce offspring at a conto the formation of sectors in range expansions with ex- 393 by dispersing in a random direction with dispersal disclusively short-ranged dispersal. However, these features 394 tances drawn from a jump kernel incorporating shortwere observed in lattice models which assumed instant lo-<sup>395</sup> ranged and long-ranged dispersal which we define below cal dynamics and founder-takes-all at the deme level; the 396 (Eq. (3)). To focus on the effects of the spreading proinfluence of slow local saturation on the evolution of het- 397 cess and enable direct comparison with previous work erozygosity could not be gauged. In this study, we aim to 398 (see Section II), our model includes two simplifying asestablish whether insights derived from lattice models of 399 sumptions. First, each individual has an allelic identity range expansions still apply in a continuous-space model 400 which is passed on exactly to offspring with no possifor which the lattice model assumptions can be violated 401 bility of new mutations; this enables us to evaluate the to a controllable degree, and to quantify the effect of ex- 402 persistence of initial neutral variation purely due to dis-



FIG. 3. An outline of the simulation procedure. a.) A snapshot of a population during a range expansion. The dots represent individuals in the population. Suppose the local carrying capacity is 5. An individual born at position A would only count three others in its local region (dashed circle centered at A), so it would survive. An individual born at position Bwould count seven others within its local region. That is too many for the individual to successfully compete against, so it would die. b.) An example jump kernel. There is a probability p of dispersing within the "local" (shaded) region, that is, within distance  $r_{\rm b}$ . The jump kernel decays according to the power law  $J(r) \sim 1/r^{\mu+1}$  beyond the local region.

405 406 407 408 409 and immobility post-establishment provide a reasonable 410 approximation for trees that produce massive numbers of 411 412 413 414 the patterns left behind by the initial expansion can still 415 be representative of long-time trends despite the subse-416 quent gene flow due to replenishment and reshuffling of 417 individuals [14, 15]. 418

In the absence of demes with a fixed carrying capacity, 419 <sup>420</sup> a different mechanism to regulate population growth is needed. We assume that the environment has uniformly-421 distributed resources which can support a uniform car-422 423 424 425 426 (Fig. 3a). The population density and interaction dis-427 428 tance can be combined to define a local carrying capacity 480 cost of added complexity and a larger parameter space.  $_{429}$  K via

$$K = \rho \pi r_{\rm b}^2. \tag{2}$$

431 432 433 434  $_{435}$  individual dies. If there are fewer than K other individu-  $_{439}$  do not use our model to draw any conclusions about the 436 als, the new individual establishes successfully in its new 490 spatial distribution of individuals on scales smaller than <sup>437</sup> location and survives for the remainder of the simulation. <sup>491</sup> the interaction distance.

The local interaction region in our continuous-space 438 simulation resembles the geographic subdivision unit (the 439 deme) used in lattice-based models. The concept of in-440 stantaneous local saturation, or a local founder-takes-441 all effect, would therefore correspond to an individual quickly filling its interaction region with its offspring be-443 fore it (or its descendants) attempted any long-range dis-<sup>445</sup> persal events. In order to smoothly depart from the as-<sup>446</sup> sumptions of the lattice model, it would be useful to con-<sup>447</sup> trol the fraction of dispersal events which are "local", i.e. <sup>448</sup> within the interaction region, as opposed to long-range. To do so, we used a two-part jump kernel that allows us 449 <sup>450</sup> to explicitly specify the probabilities of local versus long-<sup>451</sup> range dispersal, as sketched in Fig. 3b. In full, the jump 452 kernel is as follows:

$$J(r) = \begin{cases} p/r_{\rm b} & r \le r_{\rm b} \\ \frac{(1-p)}{r_{\rm b}^{-\mu}} \mu r^{-(\mu+1)} & r > r_{\rm b} \end{cases}$$
(3)

 $_{453}$  where p is the probability of dispersing within the local <sup>454</sup> region. The short-range part of the jump kernel is cho-<sup>455</sup> sen to be featureless, with the only notable property be-<sup>456</sup> ing that the integrated probability  $\int_0^{r_{\rm b}} J(r) dr = p$ . The move, die, or renew themselves. This assumption al- 457 long-range part of the jump kernel matches the power-law lows us to hone in on the dynamics of establishment 458 kernel used in the prior works discussed [16, 27, 39] and and expansion, without confounding effects or compu-  $_{459}$  the prefactor ensures the normalization  $\int_{r_{\rm b}}^{\infty} J(r) dr =$ tational expense from reshuffling and replenishment of  $_{460} 1 - p$ . Jump distances are randomly drawn from this regions that have already been saturated. Immortality 461 distribution using inverse transform sampling (detailed <sup>462</sup> procedure in SI Section VIA).

463 A few comments about our choice of jump kernel, seeds over scores of growing seasons, or perennial plants 464 Eq. (3), are in order. Our aim is not to exactly reproduce that replenish themselves in place once established. Even 465 a biologically measured jump distribution at all lengths, in populations for which these assumptions do not hold,  $_{466}$  but rather to capture the two main features of interest <sup>467</sup> in a simplified kernel—a tunable balance between short-<sup>468</sup> and long-range dispersal determined by the parameter p, 469 and a fat-tailed kernel with a specified power-law falloff  $_{470}$  controlled by the exponent  $\mu$ . For simplicity, we chose <sup>471</sup> the short-range part of the jump kernel to be constant  $_{472}$  with distance r; other forms are expected to lead to sim-<sup>473</sup> ilar results provided the integrated probability of jump  $_{474}$  lengths between 0 and  $r_{\rm b}$  evaluates to p. The chosen rying capacity per unit area, quantified by a maximum 475 form also implicitly assumes that the same length scale population density  $\rho$ . We introduce an interaction dis-<sub>476</sub>  $r_{\rm b}$  governs the interaction distance for the density regutance  $r_{\rm b}$  which demarcates a disc-shaped region within  $_{477}$  lation and the dispersal behavior. We could have built which an individual competes with others for resources 478 a model with an additional length parameter dictating 479 the spatial features of the dispersal kernel, but at the <sup>481</sup> Our simplified choice allows us to dial in a specific bal-<sup>482</sup> ance between local and long-range dispersal by adjusting  $_{483}$  the parameter p alone, which enables direct comparisons When an individual is born, it undertakes a random dis- 484 of different simulations where the kernel exponent, local persal event and counts the number of individuals within 485 carrying capacity, and size of the density regulation rethe interaction region surrounding its new location. If 486 gion are kept unchanged. Since the exact shape of the there are at least K other individuals in the interaction  $_{487}$  jump kernel at short distances is not biologically realisregion, the duplication event is unsuccessful and the new 488 tic (for instance, it has a discontinuity at  $r = r_{\rm b}$ ), we

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492 493 ronment are both featureless, and the same length scale 552 jumps out of an interaction region does depend on K. 494 495 496 497 498 all distances reported from simulations are in units of  $r_{\rm b}$ . 556 jumps; i.e.  $\tau_{\rm s} \ll \tau_{\rm j}$ . Using the above estimates for the Time units are chosen such that each individual generates 557 characteristic times, we find the criterion 499 <sup>500</sup> offspring via a Poisson process with a duplication rate of one; i.e. time is reported in units of the average gener-501 502 ation time for an individual. Note that not all offspring survive, because of the density regulation mechanism. 503

Once the length and time units have been fixed, the 504 consequential parameters are the kernel exponent  $\mu$ , the 505 probability of local dispersal p, and the local carrying 506 capacity K (which determines the local density  $\rho$  via 507 Eq. (2)). Simulations begin with 10K individuals whose 508 and y positions are random draws from a Gaussian 509 xdistribution with mean zero and standard deviation  $2r_{\rm b}$ . Everyone in the population gets a chance to produce off-511 spring every time step, which disperse according to the 512 jump kernel with relevant p and  $\mu$  and then either survive 513 <sup>514</sup> or don't depending on the population density where they happen to land. Simulations end once the population 515 size exceeds a predetermined threshold, usually four or-<sup>517</sup> ders of magnitude larger than the initial population size. See SI Section VIA for more details on the simulation 518 procedure. 519

520 521 522 523 525 526 527 our model). While the full saturation dynamics is com- 562 to second arrival will be different for a region near the 528 529 531 534  $_{535}$  fied model, the interaction region fills up according to a  $_{590}$  colonized region ( $\tau_2$  is large as well). Therefore, we use  $_{536}$  logistic function with growth rate p, for which the satu- $_{591}$  the same criterion, Eq. (4), to gauge whether both as-537 ration dynamics are set by the characteristic time scale 592 sumptions underlying the lattice model are satisfied in  $\tau_{s} \equiv 1/p$ . (The actual saturation time for a deme with  $\tau_{s}$  our continuum model. In the next section, we directly <sup>539</sup> a discrete population has an additional logarithmic de-<sup>594</sup> verify that our simulation results include regimes which 540 pendence on the carrying capacity, see SI Section VIC; 595 violate the assumptions of instantaneous local saturation 541 we ignore this weaker dependence compared to the dom- 596 (Section IV A) and local founder-takes-all (Section IV B),  $_{542}$  inant 1/p dependence in the present discussion of char-  $_{597}$  thereby departing strongly from the prior lattice models. 543 acteristic time scales.) The saturation time scale must 598 In summary, to violate the lattice model assumptions 544 545 gion to send out long-range jumps. The highest possible  $_{600}$  to or lower than 1-1/K. If we evenly sample values of 546 rate occurs when the region has saturated to population 601 p between zero and one, we find that the lattice model  $_{547}$  K and sends out long-range jumps at a rate K(1-p).  $_{602}$  assumptions are violated at most parameter values. For <sup>548</sup> Therefore, we identify  $\tau_i \equiv 1/(K(1-p))$  as the charac- <sup>603</sup> instance, if we set the carrying capacity to K = 10, the  $_{549}$  teristic time scale separating long-range jumps out of an  $_{604}$  criterion is violated for p values up to around 0.9; when

We now specify appropriate units for length and time 550 interaction region. Note that the local saturation time in our simulations. Since the individuals and the envi- 551 scale is independent of K, whereas the rate of long-range

 $r_{\rm b}$  governs both the density regulation and the disper- 553 The instantaneous local dynamics assumed in the latsal, the interaction distance is the natural length unit in 554 tice model is approached when the local saturation time our model. In our simulations, we set  $r_{\rm b}$  to one, so that 555 is much smaller than the typical time between long-range

$$\frac{p}{1-p} \gg K \tag{4}$$

<sup>558</sup> for fast local dynamics. This criterion is always satisfied 559 as  $p \to 1$ . When K is large, p must be at least 1 - 1/K<sup>560</sup> for Eq. (4) to be satisfied: for appreciable local carrying <sup>561</sup> capacities, the fraction of local dispersal events must be 562 very close to one for the criterion to hold. If an individual  $_{\rm 563}$  competes with a large number of other individuals in its <sup>564</sup> neighborhood for resources, Eq. (4) is satisfied only if the  $_{\rm 565}$  vast majority of dispersal events are local and long-range <sup>566</sup> jumps are exceedingly rare. Our estimate emphasizes the 567 need for simulations with explicit local dynamics to in- $_{\rm 568}$  vestigate the broad range of parameters where the lattice <sup>569</sup> model assumptions do not hold during jump-driven range 570 expansions.

571 The criterion  $\tau_{\rm s} \ll \tau_{\rm i}$  ensures that new migrants origi-<sup>572</sup> nate from fully saturated regions. To satisfy the second 573 assumption of the lattice model—the local founder-takes-574 all effect—we additionally require that the characteristic We next identify characteristic time scales in the prob- 575 time between the arrival of the first migrant and a polem which will enable us to choose parameters which vi- 576 tential second migrant by long-range dispersal, which we olate the instant local dynamics and local founder-takes- 577 call  $\tau_2$ , is much larger than the local saturation time scale all assumptions. (For an expanded discussion with po-  $578 \tau_s$ . Unlike  $\tau_s$  and  $\tau_i$ , however, we do not have direct contential improvements, see SI Section VIB). First let us 579 trol over  $\tau_2$ ; the latter time scale will depend not only consider the characteristic saturation time scale for a sin- 550 on the model parameters but also on the location of the gle interaction region (which takes the place of a deme in 581 region being colonized. For example, the expected time plicated because of the influence of offspring from nearby 583 core of a colony compared with a region far from the core interaction regions, we can make a simplified estimate of 584 that was recently seeded by long-range dispersal. Neverthe saturation time by considering only the descendants 555 theless, we expect that  $\tau_2$  is closely related to the time of the pioneer individual which undergo local dispersal. 556 scale  $\tau_{\rm i}$  associated with sequential long-range jumps out Assuming that all these descendants land in the same 587 of any given region: if long-range dispersal from all reinteraction region, we have an effective division rate of 588 gions is exceedingly rate ( $\tau_i$  is large), it will take a very p (in our units) for the local population. In this simpli- 589 long time for a second migrant to arrive into a newly be compared to the typical time for the interaction re- <sup>599</sup> we require local dispersal probability values comparable

8

 $_{606}$  our simulations, we choose values of carrying capacity  $K_{600}$  takes longer to fill up at low local dispersal rates and/or 607 608 <sup>609</sup> assumptions to be violated over most of these parameter <sup>603</sup> functional dependence on the carrying capacity.  $_{610}$  values, except at the upper range of values of p.

611

### IV. RESULTS

#### Local dynamics are consistent with logistic 612 Α. growth 613

We first analyze the effect of modifying the local dis-614  $_{615}$  persal probability p on the population dynamics within interaction regions. Consider the fate of the interaction region surrounding a pioneer that has landed in an empty 617 part of the range. If all local dispersal events experienced 618 by the pioneer and its offspring landed within the pio-619 neer's interaction region, we would expect exponential 620 growth of the local population with rate p until the car-621 rying capacity K is reached. In practice, the interaction 622 <sup>623</sup> regions of the offspring only partially overlap with that of <sup>624</sup> the pioneer, so the population growth levels off smoothly upon approaching the maximum value. When saturation 625 curves across many interaction regions are averaged for 626 a given set of parameters, the average curve takes on the 627 form of a logistic function as shown in Fig. 4a–b. Upon varying p and  $\mu$  independently, we find that the satura-629 tion proceeds faster as p is increased whereas it is not 630 strongly affected by the kernel exponent (Fig. 4b). 631

We use the logistic growth rate, extracted from a two 632 parameter fit to the average growth curves (see SI Sec-633 tion VIC for details), to quantify the local saturation 634 <sup>635</sup> dynamics. As expected, we find that the growth rate is largely independent of carrying capacity and is determined by the local probability p (SI Fig. 12). The 637 <sub>638</sub> growth rate remains nonzero as  $p \to 0$ , due to multi-step 639 colonization: although no direct offspring of the pioneer 640 can land in its own interaction region, the descendants <sup>641</sup> of these offspring can land within the interaction region <sup>679</sup> A nonzero heterozygosity indicates that more than one <sup>642</sup> of the pioneer which eventually gets filled. Multi-step ef- <sup>680</sup> allele is present in the region; the larger the heterozygos-643 644 carrying capacity K (plateaus at large t in Fig. 4a–b), 683 allele dominates. as outlined in SI Section VIC. The true saturation value 684 646 647 tracted from the logistic fit and is denoted as K'. 648

649 650 persal probability and not the carrying capacity, the typ-688 ties are normalized against the value at which one ex-651 652 653 655 <sup>656</sup> region. We find that the saturation time falls with in- <sup>694</sup> than one indicate that interaction regions typically have 657 creasing local dispersal levels, and rises with increasing 695 a single allele, whereas values greater than one indicate 658 local carrying capacity, as shown in Fig. 4c. The interac- 696 the expected presence of more than one allele signaling

 $_{605}$  K = 100, the criterion is satisfied only for p > 0.99. In  $_{659}$  tion region around a pioneer that seeds a distant satellite between 10 and 100, and local dispersal probabilities in 661 at high carrying capacities. Notably, the saturation time the range  $0 \le p \le 0.997$ . We expect the lattice model 662 falls linearly with p, but has a slow (roughly logarithmic)

### B. Slow local saturation invalidates 664 founder-takes-all assumption within interaction 665 regions 666

Slow saturation of the pioneer's local region increases 667 668 the chance that other individuals who are not descendants of the original pioneer will disperse into the region and establish themselves before the region is full. If an 670 671 individual who arrives later has a different allele than the <sup>672</sup> original pioneer, there will be multiple alleles within the 673 region, which introduces genetic diversity within interac-674 tion regions in stark contrast to the homogeneous demes <sup>675</sup> imposed by the lattice model. This creates a measurable <sup>676</sup> signal that local saturation times are now comparable to 677 or slower than the typical time gap between the arrivals 678 of first and second migrants by long-range dispersal,  $\tau_2$ .

To quantify the deviation of local population structure from the local founder-takes-all assumption as the saturation time is increased, we introduced neutral genetic variation in the initial population. Every individual in the initial population was assigned a unique allele, which did not affect the dispersal or reproduction dynamics but was passed on to offspring. The establishment of multiple alleles in the same interaction region was detected by computing the local heterozygosities in the interaction region of isolated pioneers. The heterozygosity, H, is the probability that any two randomly selected individuals will have different alleles. Upon counting the fraction  $f_i$ of individuals with each neutral allele i in an interaction region, the heterozygosity of that region is computed as

$$H = 1 - \sum_{i} f_i^2.$$

fects are also responsible for generating saturation curves 681 ity, the more evenly distributed the different alleles are whose final population values do not exactly equal the 682 in frequency, corresponding to a region in which no single

We averaged the local heterozygosity within the inof the population within an interaction region can be ex- 685 teraction regions of many independent pioneers to ob-<sup>686</sup> tain a characteristic measurement of the local diversity Although the logistic growth rate is set by the local dis- 687 for each parameter value. The averaged heterozygosiical time taken to fill the interaction region of a pioneer 600 pects a fully occupied interaction region to have exactly depends on both quantities. Since the logistic growth 600 one individual with a different allele than the pioneer: function is continuous and strictly reaches K' only as  $_{691}$  H<sub>N</sub>  $\equiv 2(1/K)(1-1/K)$ . With this definition, the nor- $\rightarrow \infty$ , we define the time taken to reach a local popu- 692 malized average heterozygosity  $\langle H \rangle / H_{\rm N}$  has the following lation of K' - 1 as the saturation time for an interaction <sup>693</sup> interpretation: normalized average heterozygosities less



FIG. 4. Saturation dynamics of interaction regions around pioneers. a.) and b.) show the population growth within the interaction region of pioneers (individuals which land in an empty region) as a function of time from establishment of the pioneer, averaged across many pioneers for different values of p (colors). **a.**, K = 10 and  $\mu = 1.5$ ; **b.**, K = 100 and three different kernel exponents (dashes). Each curve in panels (a.) and (b.) is the average of the local saturation around approximately 60 pioneers gathered across multiple simulations. c.) Saturation time of interaction regions, defined as the time taken for the fitted logistic growth function describing the population within an interaction region to reach one less than the saturating population (see SI Section VIC for details), for  $\mu = 1.5$ . We fit the logistic growth function to the local saturation data of approximately 60 interaction regions around pioneers and then computed the saturation time for each region based on the fitted growth rate and carrying capacity. The points in the plot are averages and the error bars are the standard deviations of the computed saturation times of individual interaction regions at each set of parameters.



FIG. 5. Influence of slow local dynamics on local diversity. Each symbol represents a triplet of parameter values  $(K, p, \mu)$ and its color shows the average heterozygosity within the interaction regions around several pioneers who seeded distant satellites, normalized against the heterozygosity at which a fully occupied interaction region is expected to have one individual with an allele different from everyone else in the region. Interaction regions are expected to be homogeneous at parameter combinations where the average normalized heterozygosity is less than one (blue points). They are expected to have more than one allele where the average normalized heterozygosity is greater than one (red points), indicating that other individuals typically disperse into and establish themselves within a pioneer's interaction region before it fills up with descendants of the pioneer. The values reported come from the averages across about 50 interaction regions gathered from multiple simulations at each set of parameters.

697 the interaction region. 698

We find that the local heterozygosity is high at low lo-690 cal dispersal rates and high carrying capacities (Fig. 5), 700 consistent with our expectations from the slow satura-701 tion dynamics in this part of parameter space. At the 702 smallest carrying capacity (K = 10), heterozygosity lev-703 704 occurs fast enough that interaction regions are filled by 714 this crossover occurs is larger for broader jump kernels 705

a deviation from the founder "taking all" at the level of 706 descendants of the pioneer individual that first arrived <sup>707</sup> in the vicinity. This situation most closely parallels the <sup>708</sup> lattice models. As the carrying capacity is increased, <sup>709</sup> however, we observe appreciable levels of heterozygosity 710 at low levels of local dispersal where the saturation dy-711 namics of regulation regions is slowest (Fig. 4c). As the 712 local dispersal rate increases, a smooth crossover occurs els are low across nearly all jump kernels: local saturation 713 from high to low heterozygosity. The value of p at which

10

715 We expect these trends to continue for carrying capacities  $\pi_1$  estimate of the core radius from our simulation data: on either side of the range we show here. For lower car-717 rying capacities, local diversity would become lower ev-718 erywhere. For higher carrying capacities, the boundary 719 between pioneer-dominated and not pioneer-dominated 720 (blue points and red points, respectively) would continue 721 to move to the right. The region of parameter space 722 where founders typically "take all" will continue to shrink 723 as carrying capacity increases. 724

In summary, measurements of local heterozygosity 725 (Fig. 5) indicate a breakdown of founder takes all over 726 wide swaths of parameter space, especially for high car-727 rying capacities and broad jump kernels. While lo-728 cal interaction regions remain largely monoallelic when 729 long-range dispersal is very rare  $(p \gtrsim 0.9)$ , we find evi-730 dence that multiple incursions into the same region leave 731 a persistent contribution to the local genetic makeup 732 within interaction regions when long-range and local dis-733 persal rates are of similar order. We next investigate 734 the extent to which these *local* deviations from founder 735 takes all impact global features of the population ex-736 pansion, and in particular whether they lead to depar-737 tures from the population-level behavior of jump-driven 738 range expansions predicted using lattice-based models in 739 Refs. [16, 27]. 740

#### С. Increased long-range dispersal favors faster 741 population growth 742

743 <sup>744</sup> der jump-driven expansions is their dramatic speedup <sup>798</sup> the two largest p values for  $\mu = 2.5$ , but is faster than 745 746 747 748 749 and are more likely to be unsuccessful due to a lack of  $_{\scriptscriptstyle 804}$ 750 751 752 753 754 755 local saturation is slower. 756

757 758 759 760 761 762 763 764 765 766 767 <sup>769</sup> cisely defining the core region. We follow Ref. 39 in using <sup>823</sup> condition from Ref. 16, Eq. (1), which is expected to

(lower  $\mu$ ): longer dispersal events favor mixing of alleles. 770 the mass-equivalent radius of the population as our best

$$\ell(t) \equiv \sqrt{\frac{M(t)}{\pi\rho}},\tag{5}$$

<sup>772</sup> which provides the required scaling  $M(t) \sim [\ell(t)]^d$  [16]. <sup>773</sup> This definition assumes that the bulk of the population <sup>774</sup> is present in regions where the population has reached its <sup>775</sup> maximum density locally. We averaged  $\ell(t)$  trajectories across different instances at each set of parameters to 777 get a *growth curve* characterizing the average growth in extent of the population. 778

We found that the acceleration of range expansion due 779 780 to long-range dispersal is preserved in the continuum <sup>781</sup> model, as shown by the growth curves in Fig. 6. We 782 focus on the behavior at long times beyond the satura-783 tion time scale of a single interaction region (which is of <sub>784</sub> order 10 for K = 10, see Fig. 4c). When all dispersal is res short-range (p = 1), the average colony size approaches 786 a linear relationship at long times (dashed curves; linear 787 fit shown with solid curves at upper right), signifying the 788 expected constant-speed outward advance of the popula-<sup>789</sup> tion front [30]. Small levels of long-range dispersal (solid <sup>790</sup> curves) are sufficient for the size to grow faster than lin-<sup>791</sup> early with time, as evidenced by a steeper slope on log-792 log axes compared to the dashed curves. The growth <sup>793</sup> at long times appears to be faster than any power law 794 (i.e. faster than linear on log-log axes) for all values of <sup>795</sup> p at  $\mu = 1.5$  and  $\mu = 2.0$  (Fig. 6a–b), in line with ex-<sup>796</sup> pectations from the lattice model. Growth approaches a The salient feature of the global population growth un- 797 power law in time with exponent greater than one at the compared to expansions that only involve short-range  $_{799}$  power-law for the smaller values of p over the population jumps: the typical radial extent of the core region  $\ell(t)$  so sizes simulated. In all cases, decreasing the probability grows faster-than-linearly with time when  $\mu < d + 1$ . <sup>801</sup> of short-range dispersal speeds up the colony expansion, This boost occurs because offspring attempting short- 802 as expected: long-range jumps are far more likely to land range jumps will land close to their parents and siblings, <sup>803</sup> in empty regions and succeed, compared to local jumps.

Many consequential features of the expansion, howlocal carrying capacity. By contrast, long-range jumps 805 ever, are determined not by the absolute growth of the tend to transport offspring to empty areas where they 806 population size with time but by the functional form establish and proliferate successfully. Therefore, lower 807 of the growth. For instance, the qualitative differences values of the local dispersal probability p are expected to  $_{808}$  in global diversity among different kernel ranges (Secfavor faster population growth overall, even though the <sup>809</sup> tion IIB) are owed to the different functional forms of  $_{s10} \ell(t)$  observed in the lattice model (see Section II for a We measured the population growth with time, M(t), summary). It would be useful to quantify whether and for many independent range expansions at each parame-<sup>812</sup> how the local dispersal rate influences the functional form ter value. To connect with the results from lattice-based \*\*\* of the population growth curves. A direct comparison models described in Section II A, we need an estimate of \$14 of the growth curves to the asymptotic forms derived the core region within which the population has reached <sup>815</sup> using the lattice model is not expected to succeed, besaturation. When growth is driven by long-range jumps, <sup>\$16</sup> cause the growth curves can take a long time to reach there is no sharp boundary between occupied and empty  $_{\$17}$  their asymptotic forms, especially for values of  $\mu$  near the regions even in the lattice model. Rather, the local den- 313 space dimension d = 2 [16]. This feature of jump-driven sity is close to  $\rho$  out to some distance from the origin, be- <sup>819</sup> growth is apparent in Fig. 6c, in which the measured vond which it crosses over to a power-law decline in den-  $_{220}$  growth curves for  $\mu = 2.5$  are nonlinear on logarithmic sity determined by the value of  $\mu$  [16, 39]. This smoothly are axes and deviate from the asymptotic power-law form varying occupancy profile leaves some ambiguity in pre- <sup>822</sup> even at long times. Instead, we use the self-consistency

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FIG. 6. Average growth curves for different  $\mu$  and p at K = 10. These are averages of growth curves from 241 individual simulations at each set of parameters. Average curves at K = 100 are shown in Fig. 11. The dashed line corresponds to simulations with only local dispersal; the resulting growth at long times is consistent with a linear relationship (gray line in the upper right of each panel).

824 hold for times beyond the local saturation time scale but <sup>825</sup> well before the time at which the asymptotic regime is <sup>826</sup> reached in the lattice model. If the entire population <sup>827</sup> in our continuum model were contained in regions that have reached local saturation at all times, then the hier-828 <sup>829</sup> archy depicted in Fig. 1 would translate to the continuum model as well, and we would expect Eq. (1) to be satisfied 830 831 exactly. This would enable us to predict future popula-832 tion growth given only the current population size and the exponent that characterizes the jump kernel. The 833 size of deviations from the exact relation could be used to quantify differences in satellite structure between the 835 continuum model and the lattice model. 836

To test the validity of the consistency condition and its 837 <sup>838</sup> ability to predict population growth, we measured the re-<sup>839</sup> lationship between the colony size  $\ell(t)$  at time t and the <sup>840</sup> quantity  $t\ell(t/2)^{2d}$  in our simulations. For t values larger than the local saturation time (order 20 or less for all pa-841 rameters, Fig. 4c), we found that the simulated growth 842 curves are consistent with a power-law relationship be-843 tween the two quantities across the entire range of local 844 dispersal probability values tested. Data for two representative values of p and two local carrying capacities are 846 <sup>847</sup> shown in Fig. 7; additional curves are shown in SI Fig. 14. For parameter values which best approximate the as-<sup>849</sup> sumption of instantaneous filling of density regulation <sup>850</sup> regions (local dispersal probability close to one and low carrying capacity), the power-law exponent quantifying 862 the time doubling hierarchy, so using Eq. (1) leads to an 851 852 853 854 855 856  $_{857}$  p or high values of K. Instead, the population size at  $_{858}$  in our continuum expansions than what is predicted by  $_{858}$  time t is larger than that predicted by the population at  $_{869}$  the time doubling hierarchy derived using lattice models  $t_{s59}$  time t/2 according to Eq. (1) (square symbols and or-  $t_{s70}$  that assume instant saturation of local regions. <sup>860</sup> ange discs in Fig. 7). The functional form of the growth <sup>871</sup> To quantify the extent of the deviation from the lattice-<sup>861</sup> curve appears to be *faster* than would be expected from <sup>872</sup> model behavior, we fit measurements of the quantity



FIG. 7. Quantitative test of the hierarchical time-doubling structure. Plots show the RHS of the consistency condition (Eq. (1)) versus the colony size  $\ell(t)$  for  $\mu = 2$ . Data are from the average of about 200 growth curves at each set of parameters, and only include the second half of the simulation to exclude expected deviations at short times (see SI Section VID for details). The scaling factor A was adjusted manually to overlay data from different p values for ease of comparison of the apparent power-law exponent (slope of curves on log-log scale). Analogous plots at  $\mu = 1.5$  and  $\mu = 2.5$  are shown in SI Fig. 14.

the expected exponent of  $d + \mu$  (compare green discs to 364 at time t/2. Note that Fig. 7 is plotted on logarithmic dashed line in Fig. 7). By contrast, the relationship no 805 axes, so the visibly slight difference between the slopes longer quantitatively matches the consistency condition 866 of sets of symbols and dashed lines corresponds to difwhen local saturation is slowed down by low values of 367 ferent power law relationships between  $\ell(t)$  and  $t\ell(t/2)^{2d}$ 

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FIG. 8. Inferring kernel exponents using the consistency condition. Symbols show the kernel exponent  $\mu_i$  inferred for the power law relationship between measured  $\ell(t)$  and  $t\ell(t/2)^{2d}$  (i.e. the slope in Fig. 7) from many individual simulations. Panels are labeled by the true kernel exponent  $\mu$  used in simulations. The dashed line indicates  $\mu_i = \mu$ . Each point represents the mean of the individual inferences from roughly 200 independent simulations and the error bars are the 95% confidence interval of the distribution of bootstrapped mean inferred kernel exponents.

 $t^{gr3} t\ell(t/2)^{2d}$  to the form  $B\ell(t)^{\nu}$  to extract the power-law ex-  $y_{12}$  tinuum model, or whether a modified version of Eq. (1) ponent  $\nu$  (see SI Section VID for details). This exponent <sub>913</sub> might be found for continuum space models. 874 was used to infer a kernel exponent  $\mu_i \equiv \nu - d$  from data such as those shown in Fig. 7, which can be compared 876 to the true kernel exponent  $\mu$ . To cover the two dis-877 tinct jump-driven growth regimes and the marginal value 914 879 880 881 882 883 885 886 887 888 889 890 891 892 893 894 ture of colony expansion as the parameter changes to 931 inferred kernel exponents in Fig. 8. 895 nearly all short-range dispersal, while anything less than 896 nearly all short-range dispersal seems to behave similarly 897 regardless of p. The deviations also systematically differ 898 depending on the local carrying capacity, with inferred <sub>932</sub> 899 900 at K = 10. 901

902  $_{903}$  ferred kernel exponent  $\mu_i$  that deviates from the true ker-  $_{936}$  cate that slow local dynamics introduce corrections to 904 905 906 908 <sup>910</sup> However, this observation by itself does not provide infor-<sup>943</sup> chy of satellite sizes determined the long-time behavior <sup>911</sup> mation about how the hierarchy breaks down in the con- <sup>944</sup> in the lattice model as summarized in Section IIB.

We can nevertheless identify the likely sources of the  $\mu = d$  separating them (as referenced in Section II), we  $_{915}$  discrepancy between the continuum and lattice models estimate the kernel exponent from growth curves of pop-<sup>916</sup> based on our knowledge of the local and global dynamulations whose jump kernels decay with  $\mu$  equal to 1.5,  $_{917}$  ics. The hierarchy in the lattice model was derived under 2.0, and 2.5. We find that the inferred kernel exponent 918 the assumption that satellites which drive the expansion is close to the true exponent when the local dispersal <sup>919</sup> originate in a core region that has reached its saturaprobability approaches one across all jump kernels and 920 tion density nearly everywhere, and whose size scales as carrying capacities tested (Fig. 8). This observation is  $_{221} [M(t)]^{1/d}$ . In our simulations, local regions take some consistent with our expectation that the limit  $p \rightarrow 1$  best  $_{922}$  finite amount of time to fill up, but they can begin sendapproximates the lattice model assumptions. However, 923 ing out long range migrants as soon as they are seeded. the inferred kernel exponent is systematically lower than 924 An appreciable fraction of satellites may be seeded by the true value for much of the range  $0 , reflecting <math>_{925}$  individuals dispersing from regions with local densities the shallower-than-expected slopes at low local dispersal  $_{226}$  between zero and  $\rho$ ; furthermore, the local density could in Fig. 7. The inferred exponent grows slowly with the 927 itself vary significantly through the population. These local dispersal probability up to  $p \approx 0.9$ , and then rises  $_{928}$  deviations become more prevalent for larger carrying casharply toward the true value as  $p \rightarrow 1$ . This suggests  $_{929}$  pacities (Fig. 4c), which would suggest larger deviations that there could be some functional change to the struc-  $_{930}$  at higher values of K consistent with the behavior of the

Altogether, measurements in the continuous-space exponents at K = 100 consistently lower than those  $_{933}$  model reveal small but consistent deviations in the pop-<sup>934</sup> ulation growth curves from the time-doubling hierarchy We have not isolated the mechanism leading to an in- 935 predicted in the lattice model. Our simulations indinel exponent  $\mu$ . The fact that  $\mu_i < \mu$  implies that the  $_{937}$  the time-doubling hierarchy over a large range of values time-doubling hierarchy from the lattice model, quanti- 938 of the local dispersal probability, consistent with our esfied in Eq. (1), does not hold exactly over much of the 939 timates of the parameter regimes for which the lattice range of p values. Furthermore, it shows again that the  $_{940}$  model assumptions break down (Section III). Next, we functional form of the population growth with time is <sup>941</sup> numerically investigate the impact of these corrections faster in the continuum model than the lattice model. <sup>942</sup> on the dynamics of global diversity, for which the hierar-

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### D. Increased local diversity boosts global 945 heterozygosity but does not overcome long-term 946 trends 947

948 <sup>949</sup> hanced local diversity generated by slow local dynamics <sup>1007</sup> mine the balance between diversification and coarsening 950 diversity. Recall the predicted evolution of heterozygos- 1009 the different growth regimes. 951 ity in prior models assuming fast local saturation and lo- 1010 952 953 954 955 956 957 959 960 961 962 963 964 965 966 968 969 971 972 of initial diversity over time to hold for other values of <sup>1031</sup> simulations. initial global heterozygosity as well. This proportion is  $^{1032}\,$ 974 975 the reported heterozygosity values by 0.5. 976

We first considered a set of parameters (K = 10,977 p = 0.5) for which each interaction region is dominated 978 by the offspring of the seed individual (Fig. 9a). This 979 situation approximates the local founder-takes-all mech-980 anism of the lattice models, but does not replicate it ex-981 actly as multiple incursions into interaction regions are 982 not strictly excluded. Despite the deviations, we find 983 that the evolution of global diversity in the continuum 984 simulations is consistent with expectations from the lat-985 tice model when different kernel exponents are compared. 986 (See SI Section VIF for a quantitative comparison.) Av-987 erage population heterozygosity has settled to a constant 988 proportion of its initial value for  $\mu = 1$ , and appears to 989 be approaching a constant value as well for  $\mu = 1.5$ . The 990 slow decay of heterozygosity for  $\mu = 2$  is expected; the 991 population may have to grow by several more orders of 992 magnitude before converging to a constant heterozygos-993 ity [27]. At  $\mu = 2.5$ , the heterozygosity decays steadily 994 with no sign of convergence to a finite value, as predicted <sup>1051</sup> 995 for lattice models in the range  $d < \mu < d + 1$ . At  $\mu = 4$ , 996 a constant heterozygosity is attained at large population 1052 997 998 tinct allelic identities (Fig.  $2\mathbf{a}-\mathbf{c}$ ). In each of the growth 1054 travel long distances and seed satellite colonies. Lattice

<sup>1002</sup> expansions), the behavior of the global heterozygosity fol-<sup>1003</sup> lows the qualitative patterns derived in the lattice model. 1004 In spite of the small quantitative differences in the hi-1005 erarchical structure of satellites merging with the core Finally, we investigated the consequences of the en- 1006 (Fig. 8), the overall differences in structure which deter-(Section IV B and Fig. 5) on the fate of the initial neutral 1008 in jump-driven expansions are maintained deep within

Next, we considered parameters K = 100, p = 0 for cal founder-takes-all effects (summarized in Section IIB): <sup>1011</sup> which the local founder-takes-all assumption is violated initial variation decayed steadily towards zero for jump<sup>1012</sup> across all kernels tested according to local heterozygosity kernels with  $2 < \mu < 3$  in two dimensions, but some pro- <sup>1013</sup> measurements. We found that the increased local diverportion of the initial diversity was preserved for broader  $^{1014}$  sity at these values (as indicated by colors in Fig. 5c)  $(\mu < 2)$  or narrower  $(\mu > 3)$  kernels. We simulated range <sup>1015</sup> contributes to higher global heterozygosities compared expansions where the initial population had equal pro- 1016 to the fast saturation region, as seen in Fig. 9b when portions of two fitness-equivalent alleles (initial global <sup>1017</sup> compared to Fig. 9a and SI Fig. 16. For instance, at heterozygosity  $H_{\rm G} = 0.5$ ) and measured the evolution of <sup>1018</sup>  $\mu = 2.5$  the heterozygosity has decayed by around 8% of global heterozygosity. While the outcome of a single sim- <sup>1019</sup> its initial value when  $M/M_0 = 10^4$  in Fig. 9b, in conulation is stochastic, we estimated the expected value of <sup>1020</sup> trast to a reduction by over 20% in Fig. 9a. The same the hetorozygosity as a function of population size by av- 1021 trend is observed at all kernel exponents: The mix of aleraging the outcomes of many independent runs at each 1022 lelic identities within each interaction region under slower set of parameters. Recall that no new mutations appear <sup>1023</sup> local dynamics provides a reservoir of genetic diversity during the expansions; here we study the long term fate 1024 that allows populations to retain much more diversity of any pre-existing diversity present in the initial pop-<sup>1025</sup> than possible under the monoallelic regions imposed by ulation rather than the emergence of some balance be- 1026 fast local saturation. Nevertheless, the qualitative trends tween the loss of diversity (e.g. due to drift) and the 1027 in diversity as the kernel exponent is varied continue to promotion of diversity due to new mutations. Although <sup>1028</sup> track the expectation for lattice models. In particular, we used a specific initial heterozygosity in our simula- 1029 the global heterozygosity steadily decays towards zero for tions, we expect the observed trends in the proportion  $^{1030}$   $\mu = 2.5$ , albeit at a slower rate compared to the K = 10

A steady decay in heterozygosity is also observed for obtained from our simulation data (Fig. 9) by dividing 1033 other values of the local dispersal probability for the same <sup>1034</sup> kernel exponent  $\mu = 2.5$ , see Fig. 9c. Slowing down local <sup>1035</sup> dynamics by increasing K and reducing p raises the value 1036 of heterozygosity at each population size, but does not <sup>1037</sup> prevent the steady decay as a function of  $M/M_0$ . These 1038 results show that at long times, the diversity-reducing 1039 effect of bottlenecks outweighs the local mixing due to 1040 slow saturation dynamics for  $\mu = 2.5$ . We expect that 1041 continual heterozygosity loss will be experienced for other 1042 kernels in the range  $d < \mu < d + 1$  as well, although the 1043 rate of decay will be very slow for kernels close to the 1044 critical value of  $\mu = d$ , and for kernels with slower lo-1045 cal dynamics (i.e. large carrying capacity and low local <sup>1046</sup> dispersal probability). In this regard, the high local het- $_{\rm 1047}$  erozygosities observed for kernels with  $\mu>2$  and low p1048 values in Fig. 5 are transients which we expect to de- $_{1049}$  cay to lower values if the expansions are allowed to run 1050 longer.

### DISCUSSION

Range expansions in populations experiencing longsizes due to the formation of persistent sectors with dis- 1053 range dispersal can be dominated by the pioneers who 1000 regimes separated by the critical kernel exponent values 1055 models that assume that these pioneers quickly saturate  $_{1001}$  of d and d+1 (2 and 3 respectively in our two-dimensional  $_{1056}$  the carrying capacity within their local interaction re-

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FIG. 9. Evolution of global heterozygosity for different kernels and local dispersal rates. a.) Average global heterozygosity as a function of the growing population size for different jump kernels, with K = 10 and p = 0.5. At these parameters, each interaction region is dominated by a single allele (Fig. 5). b.) Same as a for K = 100 and p = 0; at these parameters interaction regions tend to harbor multiple alleles. c.) Global heterozygosity curves for  $\mu = 2.5$  and different local dynamics traversing the spectrum from monoallelic to multiallelic local interaction regions (blue points to red points in Fig. 5). The K = 10 data is the same as in panel (a.) but has been truncated for this plot. Data as a function of relative population size were generated by by binning the population sizes from all available simulations and then computing the within-bin  $\langle H_G \rangle$ ; see SI Section VIE for details. Shading reports the standard error of the mean within each bin in all panels, as an estimate of the uncertainty in our estimate of the ensemble average. Data come from about 200 independent simulations for  $\mu \leq 2$  in panel (a.) and about 400 simulations for  $\mu = 2.5$  in panel (a.) and all of panels (b.) and (c.). The curve for  $\mu = 4$  in panel (a.) comes from just 24 runs since simulations with  $\mu > d+1$  take much longer and the sectoring mechanism for preserving diversity is well understood (see Fig.  $2\mathbf{a}-\mathbf{c}$ ).

1057 gion have provided many insights into the dynamics and 1089 local dispersal rates. In future work, we aim to incorpo-1059 1060 1061 1062 sal for lattice models to be accurate become increasingly 1095 migrant. strict as the local carrying capacity increases (Eq. (4)). 1096 1064 1065 1066 1067 1068 1069 1070 be valid. 1071

1072 1073 1074 1075 1076 1077 1078 1079 1080 1081 1082 1083 1084 1085 <sup>1087</sup> as they are seeded, which can occur several generations <sup>1119</sup> the repeated coarsening of diversity due to bottlenecks <sup>1088</sup> before they saturate at high carrying capacities and low <sup>1120</sup> as pioneers expand into their newly occupied surround-

population structure of such range expansions [16, 27]. 1090 rate this feature into the model of hierarchical population However, real populations operate in continuous space 1091 growth sketched in Fig. 1, which would improve the acand with local population dynamics which play out con- 1092 curacy of theoretical predictions for jump-driven range currently with the global dynamics driven by long jumps. 1093 expansions in situations where local interaction regions In particular, the limits on the rates of long-range disper- 1094 are not immediately saturated upon the arrival of a new

We investigated the effects of departing from instan-We have introduced a continuous-space simulation of 1097 taneous local saturation on both local and global mearange expansions which departs from the gridlike spa- 1098 surements of neutral diversity. Interrogating the poptial structure and instantaneous local dynamics implied 1099 ulations within individual interaction regions originally in lattice models, enabling us to quantitatively investi- 1100 seeded by a long-range dispersal event reveals that mulgate population growth and neutral diversity in parame- 1101 tiple lineages, rather than just descendants of the pioneer, ter regimes where the lattice models are not expected to 1102 become likely as local saturation becomes slower (Fig. 5): <sup>1103</sup> our continuum model violates the assumption of a strictly We found that introducing explicit local dynamics is 1104 enforced local founder-takes-all effect. Having multiple associated with slow local saturation at low local disper- 1105 lineages within interaction regions provides a reservoir sal rates and especially at high local carrying capacities 1106 of genetic diversity that also enables greater global het-(Fig. 4). By contrast, the global population growth oc- 1107 erozygosity outside the regime where local founder-takescurs faster when local dispersal rates are low, because of 1108 all applies: generically, expansions with slower local dythe increase in long-range jumps that seed satellite pop-1109 namics exhibit higher global diversity at every stage in ulations in unoccupied regions (Fig. 6). The functional 1110 the expansion (Fig. 9). Nevertheless, the enhancement forms of the population growth curves show similarities 1111 in local diversity is not sufficient to overcome long-time with those from lattice-based models (Fig. 7), but with 1112 trends in global diversity, which continue to be detersmall yet quantifiable differences (Fig. 8). We suspect 1113 mined by the kernel exponent as was shown in the latthat these differences arise due to a violation of a cen- 1114 tice model [27]. In particular, when  $\mu < 2$  the global tral assumption of the lattice model: that satellites are 1115 heterozygosity settles to a stable value after an initial seeded by long-range migrants who disperse from fully 1116 period of decay, whereas for  $2 < \mu < 3$  the heterozyoccupied source regions. In our continuum model, satel- 1117 gosity decays steadily as the range expansion progresses lites can begin sending out long-range migrants as soon 1118 albeit at a slow rate. The decay is a consequence of

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1122 1123 1124 1125 1126 lattice-based predictions is a non-trivial result in light 1185 main useful to understand the qualitative behavior of ex-1127 1128 1129 measures of population genetic variation. 1130

1131 local founder-takes-all regime was unable to detect if de- 1190 these predictions in more realistic continuum simulations. 1132 scendants of an individual other than the pioneer were <sup>1191</sup> Our results are relevant to understanding and model-1133 within a local region if they happened to have the same <sup>1192</sup> ing the dynamics of range expansions in true biological allele as the pioneer by chance. Such information would <sup>1193</sup> populations, including invasive species, populations flee-1135 be useful to investigate genealogical structure beyond the <sup>1194</sup> ing climate catastrophes, and spreading viruses. We now 1136 fate of the initial neutral diversity in the population, for 1195 have a better understanding of when the first individual 1137 example to determine if the pioneer is the most recent <sup>1196</sup> to arrive in a region of space effectively determines the 1138 common ancestor of everyone else in the interaction re- 1197 genetic outcome of all others who will later inhabit the 1139 1140 mutations during the expansion. A tool like tree se- 1199 size of the interaction region, the local carrying capac-1141 quences [41] could readily be incorporated into our com- 1200 ity, and the local dispersal probability in populations of 1142 1143 promising target of future work. Understanding the com- 1202 quantities could allow researchers to predict whether or 1144 peting effects of local and long-range dynamics on ge- 1203 not the founders will "take all" when the population ex-1145 nealogies in our forward-in-time simulations could also 1204 pands its range outwards into new territory, leading to 1146 aid the construction of backward-in-time models that in- 1205 insights about how the population will evolve. 1147 corporate long-range dispersal [42, 43]. 1148

Another promising future direction would be to incor-1149 porate ongoing local competition among all individuals 1150 in the population. In this work, we assumed that estab-1151 lished individuals never move or die, modeling popula-<sup>1207</sup> 1152 tions such as trees which release large numbers of seeds <sup>1208</sup> generate figures are available in the following GitHub 1153 annually and where young saplings stand little chance 1209 repository: https://github.com/paulose-group/explicit-1154 of outcompeting mature trees around them. However, <sup>1210</sup> local-dynamics 1155 there are many species of perennial plants, for exam-1156 ple, where younger individuals can successfully compete 1157 against older individuals in their surroundings. Incorpo-1158 rating population renewal and density-dependent com-1159 petition in simulations could provide new insights into 1212 1160 how these species evolve during range expansions. We 1213 of this paper. This work benefited from access to the 1161 suspect that such competition should accelerate the de- 1214 University of Oregon high performance computing clus-1162 cay of diversity relative to our results for  $2 < \mu < 3$  1215 ter, Talapas. 1163 (Fig. 9). Local competition can completely remove alleles 1164 from the population, whereas in our model the "losing" 1165 allele is surrounded but not lost, and retains a nonzero 1166 probability of dispersing an offspring to a faraway vacant 1167 1168 habitat.

This work provides a better understanding of the range 1169 1170 of validity and the limitations of models of long-range dispersal which rely on instantaneous saturation of local 1171 interaction regions and divide continuous space into a lat-1172 tice. We have confirmed that the conclusions of the lat-1173 tice model are upheld in populations where pioneers who 1174 disperse long distances quickly saturate their immediate 1175 surroundings with their descendants; namely in populations with low local carrying capacities and high local dis-<sup>1178</sup> persal probabilities. Even when the local founder-takes-

1121 ings (Fig. 2d-f). Our results show that this coarsening 1179 all condition is violated, we have shown that qualitative is slowed down by the increased local diversity when the 1180 trends in population growth and in the evolution of neulocal founder-takes-all assumption is violated, but it is 1181 tral diversity mirror those in the lattice model, albeit not completely mitigated and the qualitative long-term <sup>1182</sup> with measurable quantitative differences. Heuristics such trends in global diversity are similar to those predicted 1183 as the time-doubling hierarchy of Ref. 16 (Fig. 1) and the using the lattice model. This qualitative agreement with 1184 effective population of satellites identified in Ref. 27 reof recent research [35] showing that models based on a 1186 pansions under long-range dispersal in non-lattice moddiscretization of space can leave surprising artifacts in 1187 els. Researchers could employ hybrid discrete/continuous <sup>1188</sup> research strategies: identify regimes of interest using the Our method of discovering local diversity outside the 1189 heuristics of the lattice model, and then test and refine

gion or to study the accumulation of additional neutral <sup>1198</sup> same immediate area. Experimenters could estimate the putational model to study such questions, which are a <sup>1201</sup> organisms in the lab or in nature. Estimates of those

### DATA AVAILABILITY

Simulation code and the code and data necessary to

### ACKNOWLEDGMENTS

We thank Peter Ralph for helpful comments on a draft

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#### VI. SUPPLEMENTARY INFORMATION 1216

#### Simulation details 1217

1218 1219 random positions near the origin. Their x and y posi-1220 a standard deviation of  $2r_{\rm b}$ . Typically about 80% of those individuals survive the density regulation in the 1222 1223 periodic boundary conditions have no effect. 1224

1225 1226 bility of mutations, so offspring have the same allele as 1227 their parent. Dispersal distances are drawn using inverse <sup>1264</sup> left, a situation termed "scramble competition" in ecol-1228 transform sampling. Recall, the jump kernel is

$$J(r) = \begin{cases} p/r_{\rm b} & r \le r_{\rm b} \\ \frac{(1-p)}{r_{\rm b}^{-\mu}} \mu r^{-(\mu+1)} & r > r_{\rm b} \end{cases}$$
(6)

 $_{1229}$  where  $r_{\rm b}$  is the boundary between local and long-range  $_{1230}$  and p is the probability of dispersing within the local region. 1231

We begin the sampling procedure by drawing a random 1232 1233 number X from the uniform distribution between 0 and 1234 1. That number X is taken to be the probability of 1235 drawing a dispersal distance less than or equal to some <sup>1277</sup> orders of magnitude, so simulations were ended once the 1236 distance x (i.e. the integral of the jump kernel from 0 to 1278 populations exceeded  $10^6$  or  $10^7$  individuals for K equal  $_{1237}$  x). Solving for x gives us our dispersal distance.

need to consider the first term of the jump kernel.

$$X = \int_0^x p/r_{\rm b} \, dr = \frac{px}{r_{\rm b}} \to x = \frac{Xr_{\rm b}}{p}$$

If X > p, the offspring disperses a long distance, so we 1286 function. have

$$X = p + \frac{(1-p)\mu}{r_{\rm b}^{-\mu}} \int_{r_{\rm b}}^{x} r^{-(\mu+1)} dr$$

which leads to

$$x = \left(\frac{(1-p)r_{\rm b}^{\mu}}{1-X}\right)^{1/\mu}$$

1238 form distribution between 0 and  $2\pi$ . 1239

1240 1241 duce offspring each time step. Offspring generation is 1296 attempts is one. The rate of divisions that land within 1242 1243 ber of offspring for each individual is a random draw 1298 the logistic growth. Therefore, the characteristic satura-1244 from the Poisson distribution with mean 1. Then all 1299 tion time of local regions is  $\alpha = p$ , and as a zeroth-order 1245 newborns simultaneously count how many other individ-1300 estimate we have  $\tau_{\rm s} \approx 1/p$ . This form is only useful for  $_{1246}$  uals are within their density regulation regions. This  $_{1301}$  p close to one, because it ignores the effect of secondary 1247 means newborns will count other newborns if they hap-1302 events which land in the interaction region. As a result, 1248 pen to land near each other by chance. It also means that 1303 the true dependence of  $\alpha$  on p is weaker:  $\alpha$  grows from 1249 space often doesn't quite fill up to the population den-1304 0.4 to 1 as p varies from zero to one (Fig. 12a). Therefore  $\mu_{250}$  sity  $\rho = K/(\pi r_{\rm b}^2)$  (as defined in section III). SI Fig. 10 1305  $\tau_{\rm s}$  varies weakly from roughly 2.5 to one over the range

<sup>1251</sup> shows how regions of space may appear habitable, and 1252 indeed would be if one single offspring were generated 1253 and counted its neighbors at a time, but do not saturate 1254 since everyone in the population typically produces an Simulations begin with 10K individuals who are given <sup>1255</sup> offspring every generation and all newborns count their 1256 neighbors simultaneously. All individuals produce one tions are random draws from a normal distribution with <sup>1257</sup> offspring per generation on average, so a region saturated  $_{1258}$  to K individuals is expected to have roughly K newborn 1259 individuals attempting to establish within that same refirst time step. The spatial landscape is large, so the <sup>1260</sup> gion every generation. Our density regulation mechanism <sup>1261</sup> mimics the biological scenario where none of those new-Offspring are produced by cloning without the possi-<sup>1262</sup> born individuals are able to get enough resources to survive since there are so many competing for what little is <sup>1265</sup> ogy [44]. The typical population that is actually attained 1266 in a local density region, which we term K', is estimated <sup>1267</sup> using a fit to a logistic growth curve (see SI Section VIC 1268 below), and deviates by at most 20% from K (SI Fig. 12). <sup>1269</sup> Alternative choices for the density regulation step, such 1270 as randomly choosing a subset of newborns to survive so 1271 that local density regions can saturate up to the target  $_{1272}$  value K (the "contest competition" scenario), could also 1273 be implemented, but at the cost of additional computa-1274 tional resources which would affect the maximum population sizes and growth times that could be simulated. 1275

We typically let the populations grow by about four 1276 1279 to 10 or 100, respectively. This allowed core radii to If  $X \leq p$ , the offspring disperses locally, so we only <sup>1280</sup> grow by about two orders of magnitude, as shown in the <sup>1281</sup> average growth curves at K = 10 in Fig. 6 and at K =1282 100 in SI Fig. 11. The solid line indicating the linear 1283 relationship between  $\langle \ell(t) \rangle$  and t at p = 1 in Fig. 6 was <sup>1284</sup> generated by fitting a line to the average growth curve 1285 from generations 100 to 1000 using NumPy's polyfit()

#### В. Time scales

The assumption of instant local saturation in the lat-1288 tice models relied on a separation of time scales between 1290 local and global dynamics: it is valid provided the time 1291 scale for saturation of local regions  $\tau_{\rm s} = 1/\alpha$  is small com-<sup>1292</sup> pared to the typical time between long-range dispersal The dispersal direction is chosen at random from the uni-1293 attempts from each "deme" or interaction region, which <sup>1294</sup> we call  $\tau_i$ . In our tunable model, our time units are set All individuals in the population get a chance to pro- 1295 such that the characteristic time between reproduction the first thing that happens each time step; the num- 1297 the interaction region is p, which sets the time scale of

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FIG. 10. Space doesn't quite fill up to the local maximum population density  $\rho$ . The two panels are snapshots from simulations that grew to roughly fifty thousand individuals. The blue points are individuals, and the color of the small yellow and green dots represents how many individuals are within a distance of  $r_{\rm b}$  of that point. Yellow points represent saturated regions; an individual born there would count at least K within its density regulation region. Non-yellow points look hospitable, and one individual born there would count less than K within its density regulation region, but individuals can't fill those spaces since typically everyone produces an offspring every generation and the newborns "destructively interfere".



FIG. 11. Average growth curves for different  $\mu$  and p at K = 100. Average curves for K = 10 are shown in Fig. 6. These are averages of about 140 growth curves at each set of parameters.

 $_{1306}$  of p values. For a more accurate estimate, we can use the  $_{1312}$ <sup>1307</sup> phenomenological form  $\alpha \approx (1+p)/2 \Rightarrow \tau_{\rm s} \approx 2/(1+p)$  <sup>1313</sup> 1308 which does not diverge as  $p \to 0$ .

When regions have reached local saturation, the rate at which each interaction region sends out long-range jumps by jumps out of regions that have reached saturation, we reduces to

$$\frac{1}{K(1-p)} \gg \frac{1}{p},$$

or

$$K \ll \frac{p}{1-p}.$$

1309 1310 explicitly do not satisfy the separation of time scales as-<sup>1311</sup> sumed in the lattice model.

### С. Logistic growth description of population dynamics within interaction regions

1314 We started the logistic growth measurements by is K(1-p). If we assume that the expansion is driven  $\frac{1}{1315}$  searching for sufficiently isolated individuals. To find 1316 individuals worthy of tracking, we searched at the end have  $\tau_{j} = 1/(K(1-p))$ . Therefore, the condition  $\tau_{j} \gg \tau_{s}$  is represented by the provided of the pro <sup>1318</sup> within a distance of  $10r_{\rm b}$ . Those individuals must have <sup>1319</sup> dispersed a long distance. We searched at the end of ev-1320 ery generation until we found at least a predetermined 1321 minimum number of isolated individuals at the same <sup>1322</sup> time. We required several at the same time purely for 1323 convenience on the data processing side; these measure-1324 ments could just as well have been gathered one at a <sup>1325</sup> time as we found the isolated individuals. Nevertheless, 1326 once we found the isolated individuals, we recorded the According to this criterion, most of our simulations 1327 number within their interaction regions every generation <sup>1329</sup> was then used to fit for logistic growth parameters. We



FIG. 12. Fitted logistic growth parameters. a.) The growth rate increases with increasing short-range dispersal as expected. It does not depend on the carrying capacity because the growth rate is determined by the early growth of the population before the density regulation restricts population growth. **b.**) The fitted local carrying capacity increases slightly with increasing short-range dispersal. Regions often don't saturate all the way to K as discussed above and shown in Fig. 10. We fit the logistic growth function to the saturation data of about 60 interaction regions across multiple simulations at each set of parameters. The points are averages and the error bars are standard deviations of the individual fits. This data comes from expansions with  $\mu = 1.5$  and is what formed the saturation times reported in Fig. 4c.

1362

1330 fit to the logistic function of the form

$$N(t) = \frac{K'}{1 + \left(\frac{K' - N_0}{N_0}\right)e^{-\alpha t}}$$
(7)

where N(t) is the population at time t, K' is the local 1331 <sup>1332</sup> carrying capacity,  $\alpha$  is the growth rate, and the initial population is  $N_0 = 1$ . We used SciPy's curve\_fit() func-1333 <sup>1334</sup> tion to make the fits and obtain K' and  $\alpha$ . We performed 1335 the fits on all individual interaction regions around the initially isolated individuals that we found that filled up 1337 values and standard deviations are shown in SI Fig. 12. 1338 <sup>1340</sup> region by setting the population size in Eq. (7) equal to <sup>1371</sup> in Fig. 8. Colony growth remains self-consistent once <sup>1341</sup> K' - 1 and then solving for t, which leads to

$$t_{\rm sat} = \frac{1}{\alpha} \log\left( (K' - 1)^2 \right) \tag{8}$$

 $_{1343} \alpha$  is itself proportional to p (see SI Section VIB), we  $_{1378}$  of data points after the first one to check for agreement <sup>1344</sup> find a logarithmic dependence of the saturation time on <sup>1379</sup> with the consistency condition and to estimate the kernel 1345 the local carrying capacity, which arises from the discrete 1380 exponent. Our run times were just enough at the lowest 1347 <sup>1348</sup> region for which we fit the logistic growth function, using <sup>1383</sup> data at  $\mu = 2$  against Eq. (1) in Fig. 7; analogous plots values of  $\alpha$  and K' from the fits to the logistic function. 1384 at  $\mu = 1.5$  and 2.5 are shown in SI Fig. 14. 1349 We report averages and standard deviations at  $\mu = 1.5_{1385}$ 1350 in Fig. 4c. 1351

1352 <sup>1353</sup> dividual in the initial population had a unique allele. <sup>1388</sup> local dispersal. Expansions at high local dispersal take 1354 1355 <sup>1356</sup> growth as described above (i.e. initially isolated individ-<sup>1391</sup> parent in regions that may already be saturated, which 1357 uals). The heterozygosities reported in Fig. 5 are aver- 1392 means the times t and sizes  $\ell(t/2)$  used here are higher ages of heterozygosities measured across typically about 1393 at high local dispersal. A scaling factor of  $A \approx 20$  was

1360 simulations and gathered from initially isolated individ-<sup>1361</sup> uals in multiple different simulations.

### Quantitative assessment of time-doubling D. hierarchy

We assessed the validity of simulation run times us-1364 <sup>1365</sup> ing the consistency condition  $\ell(t)^{d+\mu} \sim t\ell(t/2)^{2d}$  (eq. 1). <sup>1366</sup> The consistency condition is only valid after enough time <sup>1367</sup> has elapsed for long dispersal distances to be the driving to at least 60% of the local carrying capacity K. Average <sup>1368</sup> factor behind a colony's growth [16]. It is necessary to <sup>1369</sup> avoid the early times when applying the consistency con-We computed the saturation time for an interaction <sup>1370</sup> dition, such as when estimating the kernel exponent as 1372 the consistency condition becomes valid. For simulations 1373 that ran for T time steps, the values of t we used when 1374 applying the consistency condition ran from T/2 to T, 1375 so the values of t/2 ran from T/4 to T/2. The first data 1376 point we used with the consistency condition is marked  $_{1342}$  In addition to the dominant dependence ~  $1/\alpha$ , where  $_{1377}$  with a red × in SI Fig. 13. We need at least a handful nature of the local population within a deme. We com- 1381 probabilities of local dispersal and gave us many useful puted the saturation time for every individual interaction 1382 data points at high local dispersal. We compared our

The expansions at high local dispersal require many 1386 more time steps to reach the predetermined population For the local heterozygosity measurements, every in- 1387 threshold necessary to end simulations than those at low We tracked the heterozygosity in the interaction regions 1389 longer and grow slower than those at low local dispersal of the same individuals for whom we measured logistic 1390 since offspring are much more likely to land near their 1359 50 separate interaction regions in the final generation of 1394 needed to raise the points at low local dispersal to the



FIG. 13. We show  $t\ell(t/2)^{2d}$  plotted against  $\ell(t)$  from average growth curves at all sets of parameters that went into figure 7. The dots are points from all even-numbered time steps. Simulations with high values of p require many more time steps to reach a given population size than those with low values of p. We use only the points after the red  $\times$  in figure 7 and for inferring the kernel exponent as in figure 8. For simulations to be "long enough," we needed at least a handful of points once the growth became self-consistent (i.e. linear on these plots). Using average growth curves from expansions at intermediate probabilities of local dispersal result in plots somewhere between these two extremes: more data points in the linear sections than the p = 0case but not as many as in the p = 0.99 case.



FIG. 14. Analogous to Fig. 7.

<sup>1395</sup> level of those at high local dispersal. Bringing them to-<sup>1404</sup> kernel exponent by performing this fit using data from 1396 gether highlights the difference in power laws (slopes) 1405 only a later subset of the time steps as discussed in the between the sets of points at each value of K. 1397

using SciPy's curve\_fit() function. We obtained values the parameters) to get the averages and confidence intervals the for both the prefer to the Parameters in Fig. 6. <sup>1401</sup> for both the prefactor B and the exponent  $\nu$ , but only

1406 previous paragraph. We then compute the inferred ex-<sup>1407</sup> ponent as  $\mu_i \equiv \nu - d$ . We computed  $\mu_i$  using all available growth curves by fitting  $B\ell(t)^{\nu}$  to the quantity  $t\ell(t/2)^{2d}$  are growth curves (typically about 200 at any given set of using  $\Omega$  in  $\Omega$  is a set of using  $\Omega$  is a set of \Omega is a set of using  $\Omega$  is a set of \Omega is a set of using  $\Omega$  is a set of \Omega is a set of \Omega is a set o

 $_{1402}$  the exponent was of any interest for estimating the ker- $_{1411}$  The exact value of  $\mu_i$  somewhat depends on the fit 1403 nel exponent from the growth curves. We estimated the 1412 method. For comparison, we repeated the process of ex-

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1487

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power law fit, lin. space Inferred kernel exponent  $\mu_{
m i}$ 1.45 linear fit, log-log space 1.40 1.35 1.30 1.25 1.20 1.15 0.0 0.2 0.4 0.6 0.8 1.0 Probability of local dispersal p

FIG. 15. Example comparison between inferring kernel exponents by directly fitting power laws in linear space and fitting  $^{1457}$  functions of time. for the slope in log-log space. This example data comes from 1458 independent simulations below p = 0.5 and roughly 150 simulations above p = 0.5. The error bars are the 95% confidence interval of the distribution of bootstrapped mean inferred kernel exponents.

data in log-log space, where the exponent could be found <sup>1469</sup> points for Fig. 9 and SI Fig. 16. from the inferred slope. These two values would exactly  $^{\rm 1470}$ 1415 match if we had infinitely long simulations that had per- $^{1471}$  uncertainty in  $\langle H_{\rm G} \rangle$ . A consequence of the binning pro-1416 fectly converged to constant power laws, but in practice <sup>1472</sup> cedure is that standard errors of the mean heterozygosity 1417 that is not the case. There is often a slight difference <sup>1473</sup> get vanishingly small in Fig. 9 and SI Fig. 16, despite the 1418 between the average values of  $\mu_i$  from the two proce-<sup>1474</sup> fact that the heterozygosity trajectory can vary quite a 1419 dures, as shown at the example parameters in SI Fig. 15. <sup>1475</sup> bit between individual simulations. The bins in those 1420 However, we take the generally overlapping error bars as <sup>1476</sup> figures often consist of multiple data points from each 1421 a signal that it's safe to proceed with our inferred val-<sup>1477</sup> simulation, especially for the bins at larger population 1422 ues. This sort of comparison could be used as a test of <sup>1478</sup> sizes. Even though we generally have 200-400 indepen-1423 whether or not population growth has converged to the 1479 dent simulations at each parameter combination shown 1424 expected time-doubling hierarchy: consistent gaps be-<sup>1480</sup> in those figures, the points in the figures are often aver-1425 tween error bars are a warning that simulations may not 1481 ages of thousands of data points that fall within each bin, 1426 be long enough. This test led us to run longer simula-1427 tions to generate the data shown at  $p \le 0.5$  in SI Fig. 15 <sup>1483</sup> standard error of the mean is  $s/\sqrt{N}$  where s is the sam-1428 and the corresponding data points in Fig. 8. The longer <sup>1484</sup> ple standard deviation and N is the number of samples. simulations ran until they reached population sizes of 30<sup>1485</sup> Such a large number of samples within each bin gives us 1430 <sup>1431</sup> million individuals, triple the size of our usual cutoff for <sup>1486</sup> a small uncertainty in our estimate of the average  $\langle H_{\rm G} \rangle$ . 1432 simulations with K = 100.

### Reporting the evolution of global heterozygosity 1433 **E.**

1434 display the global heterozygosity as a function of the 1490 rameter values K = 10, p = 0.5 where the averaged nor-1435 growing population size as in Fig. 9 and SI Fig. 16; here 1491 mal heterozygosity is less than one (see Fig. 5), approx-1436 we discuss some options and justify our choice. The true 1492 imating local founder-takes-all, and lattice-based simu-1437 independent variable in our simulations is time. Every 1493 lations where founder-takes-all is imposed at the deme 1439 time step consists of offspring generation and dispersal 1494 level. The initial conditions in the two types of simula- $_{1440}$  followed by density regulation as discussed in Section III  $_{1495}$  tions were not exactly matched: both began with a 50/50

<sup>1441</sup> and SI Section VIA. Population size and heterozygosity are recorded at the end of each time step, after individuals have been removed from the population if their 1443 birthplaces are too densely occupied. This suggests that 1444 the "ground truth" for reporting the evolution of global heterozygosity might be plots of  $\langle H_{\rm G} \rangle$  versus time, where 1446 averages and standard errors are computed with all avail-1447 able data at a given time step. 1448

However, for generalizing results or comparing with the 1449 results of Ref. 27, it would be useful to compute  $\langle H_{\rm G} \rangle$  as 1450 a function of the population size. One method of doing 1451 this would be to compute the averages  $\langle H_{\rm G} \rangle$  and  $\langle M/M_0 \rangle$ 1452 each time step. This method ignores what can be signif-1453 icant variation in population growth rates between indi-1454 1455 vidual simulations and generates points whose horizontal <sup>1456</sup> and vertical coordinates in the plots of Fig. 9 are both

We sought to compute  $\langle H_{\rm G} \rangle$  directly as a function growth curves with K = 100 and  $\mu = 1.5$  (same data as 1459 of population size by generating binned population sizes the orange points in the left panel of Fig. 8); analogous plots 1460 and computing the average heterozygosity from all availlook similar at other pairs of parameters. Each point repre-1461 able simulation time steps where the population size was sents the mean of the individual inferences from roughly 50  $_{1462}$  within a given bin. This means that a single simulation <sup>1463</sup> can contribute to a given data point on the  $\langle H_{\rm G} \rangle$  versus  $_{1464}$   $M/M_0$  curve multiple times or not at all depending on 1465 how many time steps the population size was within that <sup>1466</sup> bin in that simulation. We used the R function cut() to <sup>1467</sup> place population sizes within 20 bins of equivalent width <sup>1413</sup> tracting  $\mu_i$  by finding the best fit line to the relevant <sup>1468</sup> in logarithmic space, thus generating equally spaced data

We use the standard error of the mean to estimate our

### Direct comparison to lattice model F.

We used data for lattice-based simulations from Ref. 27 1488 There are multiple reasonable ways to compute and 1489 to compare results between continuum simulations at pa-

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the continuum simulations with fast local dynamics (K = 10,p = 0.5; solid lines and shading are same as in Fig. 9a) to that of the lattice-based simulations reported in Ref. 27 (dashed lines). In both cases, the initial population had a 50/50 mix of two alleles (initial heterozygosity of 0.5). Kernel exponents same color refers to  $\mu = 2.4$  (lattice).

1496 mix of two alleles, but the continuum simulations began 1529 population sizes is higher than that from lattice-based 1497 1498 1499 1501 1502 1503 1504 generate data at  $\mu = 2.5$ , so we include their data from 1537 of the lattice model.

1505  $\mu = 2.4$  as a comparison with our  $\mu = 2.5$  data.

We observe that the difference in initial conditions leads to different dynamics at early times/small popu-1507 1508 lation sizes. In the continuum simulations, most of the 1509 early dynamics involves local events which mix and even <sup>1510</sup> out the starting population near the origin, and signifi-1511 cant changes in heterozygosity only kick in when the population has reached ten times its initial size. By contrast, 1512 the lattice simulations only included long-range jumps, 1513 and the heterozygosity begins to fall earlier. This dis-1514 crepancy leads to early differences in the observed heterozygosities between the two sets of models. However, 1516 <sup>1517</sup> the later trends, especially the contrast between a quick saturation of heterozygosity to a constant value at  $\mu = 1$ <sup>1519</sup> as opposed to a persistent decay for  $\mu = 2.5$  and a decay FIG. 16. Direct comparison of heterozygosity evolution in  $_{1520}$  followed by a delayed saturation for  $\mu = 4.0$ , are success-<sup>1521</sup> fully captured by the lattice model. The quantitative 1522 discrepancy between the lattice and continuum values of  $_{1523}$   $\langle H_{\rm G} \rangle$  is largest at  $\mu = d = 2$ , which is a special point <sup>1524</sup> for the underlying dynamics that leads to extremely slow match at all values except  $\mu = 2.5$  (continuum), for which the <sup>1525</sup> changes in the heterozygosity [16, 27]; we hypothesize 1526 that the small discrepancy in the initial conditions per-<sup>1527</sup> sists the longest at this special kernel exponent.

We also observe that the continuum  $\langle H_{\rm C} \rangle$  at large 1528 with typically about 80 individuals near the origin (SI 1530 simulations for all jump-driven kernels ( $\mu < 3$ ). This Section VIA) while the lattice-based simulations began 1531 is consistent with the observation that while local hetwith 111 occupied demes packed in a disc around the 1532 erozygosity is small in the continuum simulations, it is origin. Note that a deme is roughly a discrete analogue 1533 not zero for the chosen parameter values of K = 10, of an interaction region, so the continuum simulations'  $_{1534} p = 0.5$  (left panel in Fig. 5) and the slight deviations  $\overline{80}$  individuals correspond to roughly 80/K = 8 occu- 1535 from local founder takes all promote higher heterozygospied demes. Another discrepancy is that Ref. 27 did not 1536 ity compared to the strict founder-takes-all assumption

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